

# Post-processing of COSMIC open-loop radio occultation signals at CDAAC

Sokolovskiy, S., C. Rocken, W. Schreiner, D. Hunt, and J. Johnson

University Corporation for Atmospheric Research

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**1995:** Problems of PLL in moist LT found

OL tracking (raw sampling) applied for planetary occultations used high sampling rate (50 kHz)

**2001:** Principles of model-aided RO OL tracking developed

- sampling RO signal at minimal (Nyquist) rate;
- reconstruction of RO signal from complex samples

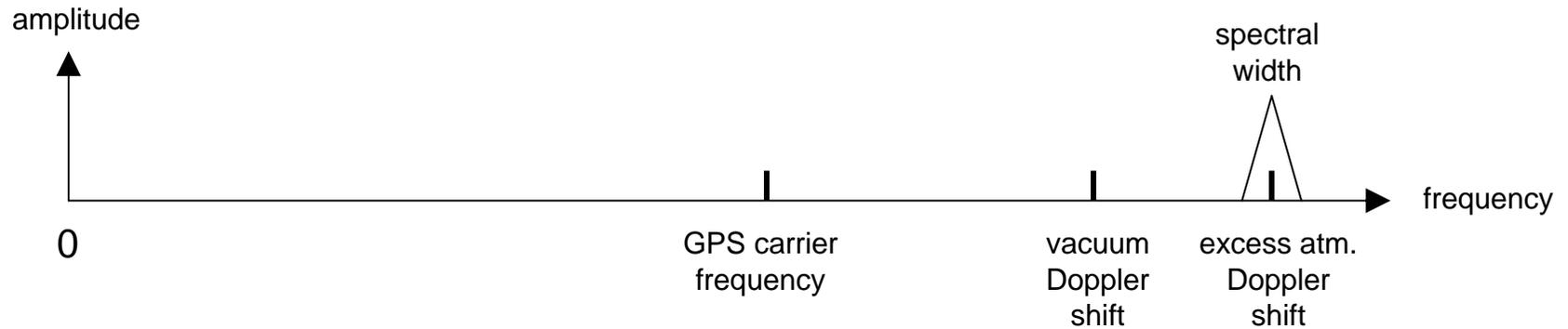
**2005:** OL tracking of GPS RO L1 signal tested on SAC-C by Jet Propulsion Laboratory

**2006:** OL tracking applied routinely on six COSMIC satellites

**This talk:** Summary of the principles and their validation based on reconstruction of COSMIC OL signals

- validation of the two frequency models used in the OL PP;
- validation of the two methods of removal of the NDM

## Spectrum of RO signal



Minimal sampling rate is limited by the spectral width

Raw sampling at minimal rate:

Aliasing of the signal spectrum + aliasing of noise = low SNR

**Needed:** noise filtering before sampling; reduction of frequency; modeling of the frequency

**Model-aided sampling rates:**

modeling of carrier frequency ~ 50 kHz

modeling of Doppler shift in a vacuum ~ 3 kHz

modeling of excess Doppler shift in the atmosphere ~ 50 Hz

## In receiver:

1) **down-conversion** (the phase model does not use feedback)

$$u_{down}(t) = u_{in}(t) \exp[-i\Phi_{rec\ mod}(t)]$$

2) **low-pass filtering** (integration)

$$I_{out}(t) = \Delta t^{-1} \int_t^{t+\Delta t} \text{Re}[u_{down}(t')] dt'$$

$$Q_{out}(t) = \Delta t^{-1} \int_t^{t+\Delta t} \text{Im}[u_{down}(t')] dt'$$

$$\Delta\Phi_{out} = \arctan 2(Q_{out}, I_{out})$$

3) **output signal sampling**

$$u(t) = A(t) \exp[i\Phi_{rec\ mod}(t) + \Delta\Phi_{out}(t)] = I(t) + iQ(t)$$

**Important:** output signal contains NDM; sampling intervals aligned with received NDM bit sequence, not with standard GPS time tags

**Important:** the only purpose of receiver model is noise filtering.

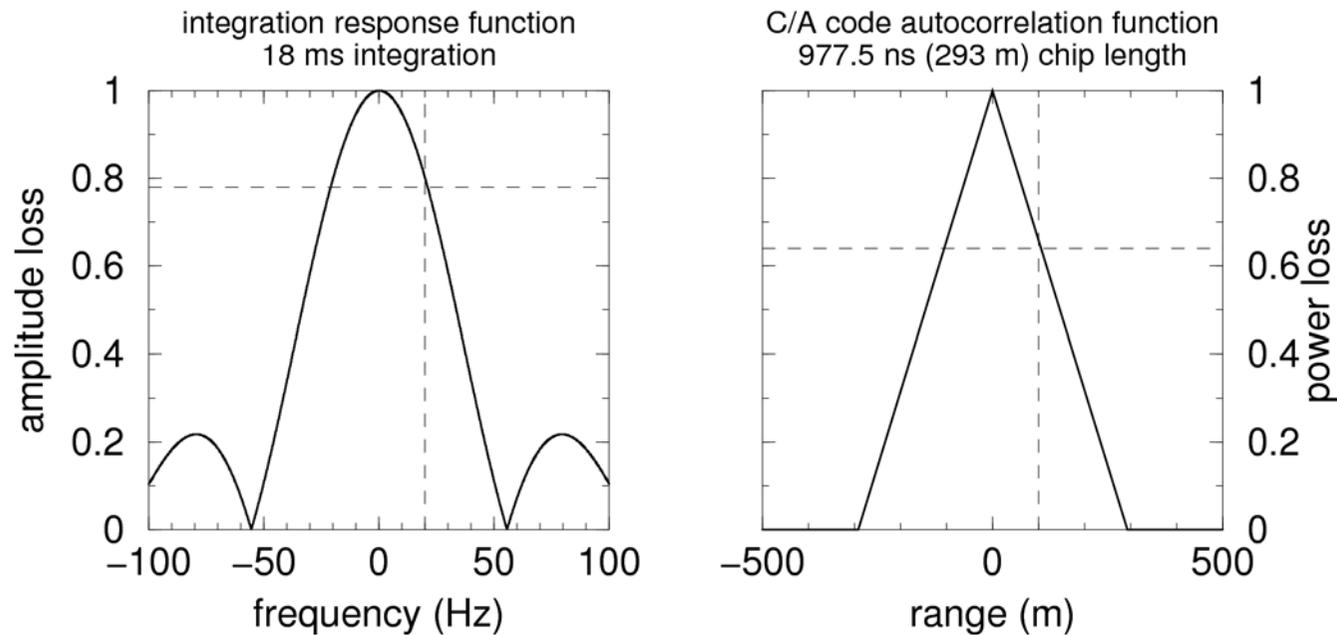
The phase is connected between samples in the post-processing.

**In receiver:** two models run in real time without a feedback

Phase (frequency) model – frequency reduction for noise filtering

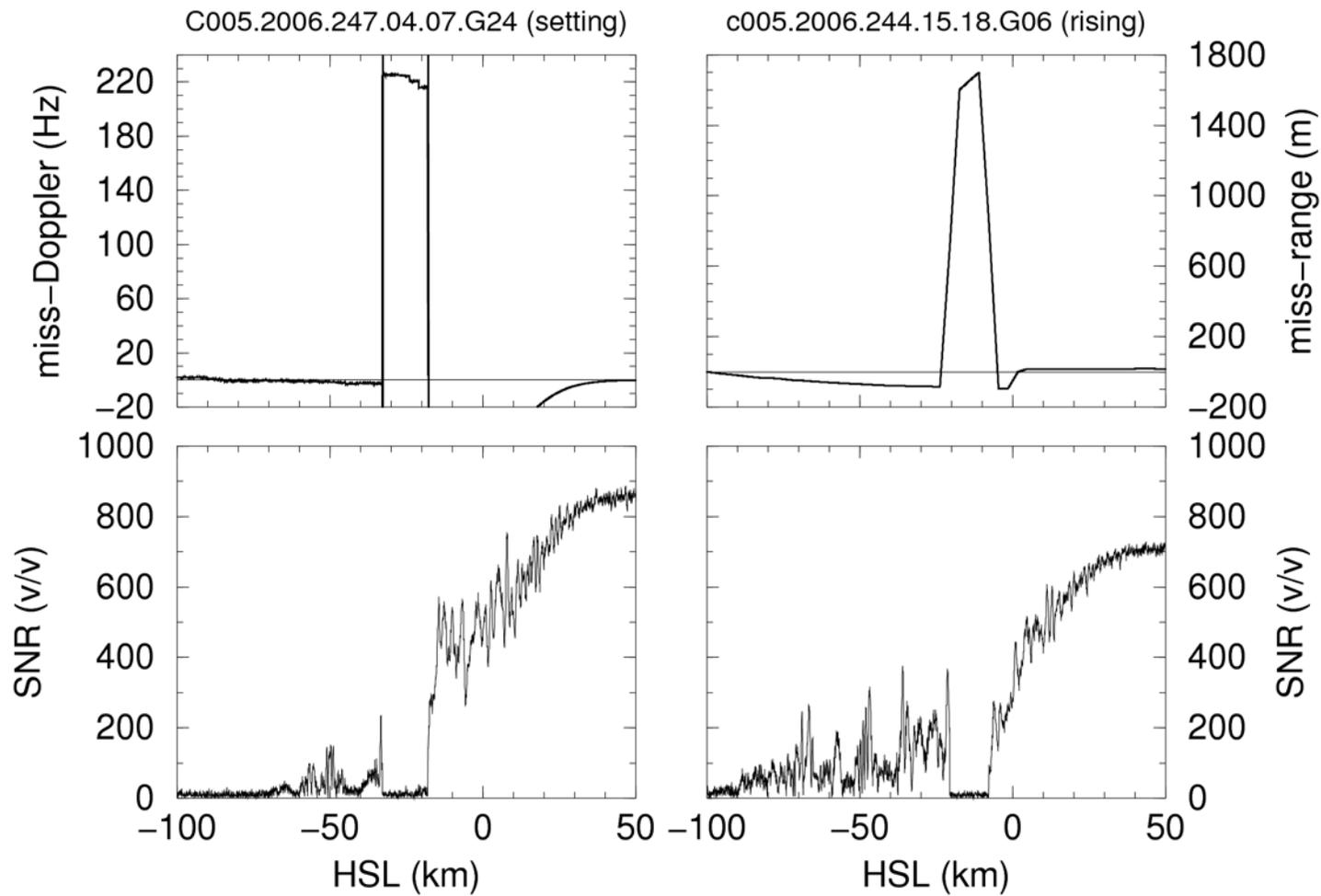
Range model – C/A code demodulation

Miss-modeling results in reduction of SNR (OL SNR lower than PLL)



# Examples of large frequency and range miss-modeling

HSL – height of straight line between GPS and LEO



In the post-processing: OL RO signal is considered as the sequence of complex samples with un-connected phase.

**Important:** receiver model is not used in the post-processing

1) **down-conversion** (frequency model discussed later)

$$u_{down}(t) = u(t) \exp[-i\Phi_{mod}(t)]$$

2) **removal of the NDM** (discussed later)

3) **connection of the phase between samples**

$$\Delta\Phi = \arctan 2(Q, I)$$

$$\Delta\Phi_i = \Delta\Phi_i + 0 \quad \text{or} \quad +2\pi \quad \text{or} \quad -2\pi$$

whichever minimizes  $|\Delta\Phi_i - \Delta\Phi_{i-1}|$

$$\Phi = \Phi_{mod} + \Delta\Phi$$

**Frequency model 0** – based on the post-processed orbits and clock and the bending angle model (impact height, latitude, month).

The bending angle model is calculated from refractivity model based on CIRA+Q climatology [Kirchengast et al., 1999].

- 1) given: radii of the GPS and LEO;  
central angle between the GPS and LEO;  
bending angle (impact height);  
solve for: zenith angles of a ray at GPS and LEO
  
- 2) given: zenith angles of a ray at GPS and LEO;  
velocities of the GPS and LEO  
calculate: frequency

**Frequency model 1** – Frequency model 0 adjusted by use of the feedback from RO signal reconstructed with the use of the model 0

**Internal removal of NDM** (similar to that in PLL) requires phase lapse between samples (atmospheric modulation + residual frequency + noise) < 1/4 cycle

If  $(I_i I_{i-1} + Q_i Q_{i-1}) < 0$  then  $I_i = -I_{i-1}$  and  $Q_i = -Q_{i-1}$

- results in 2-quadrant phase extraction;
- works well for high-latitude occultations;
- results in errors for tropical occultations in LT
- these errors contribute to unstable operation of PLL in LT

**External removal of NDM** by aligning the externally recorded NDM bit sequence with RO signal samples by accounting for the propagation time between GPS and LEO.

- results in 4-quadrant phase extraction

COSMIC deployed a network of the ground-based receivers for collection of the NDM bit sequences (USA (Boulder), Germany (Potsdam), Taiwan, New Zealand, South Africa, Brazil).

## High-latitude occultation

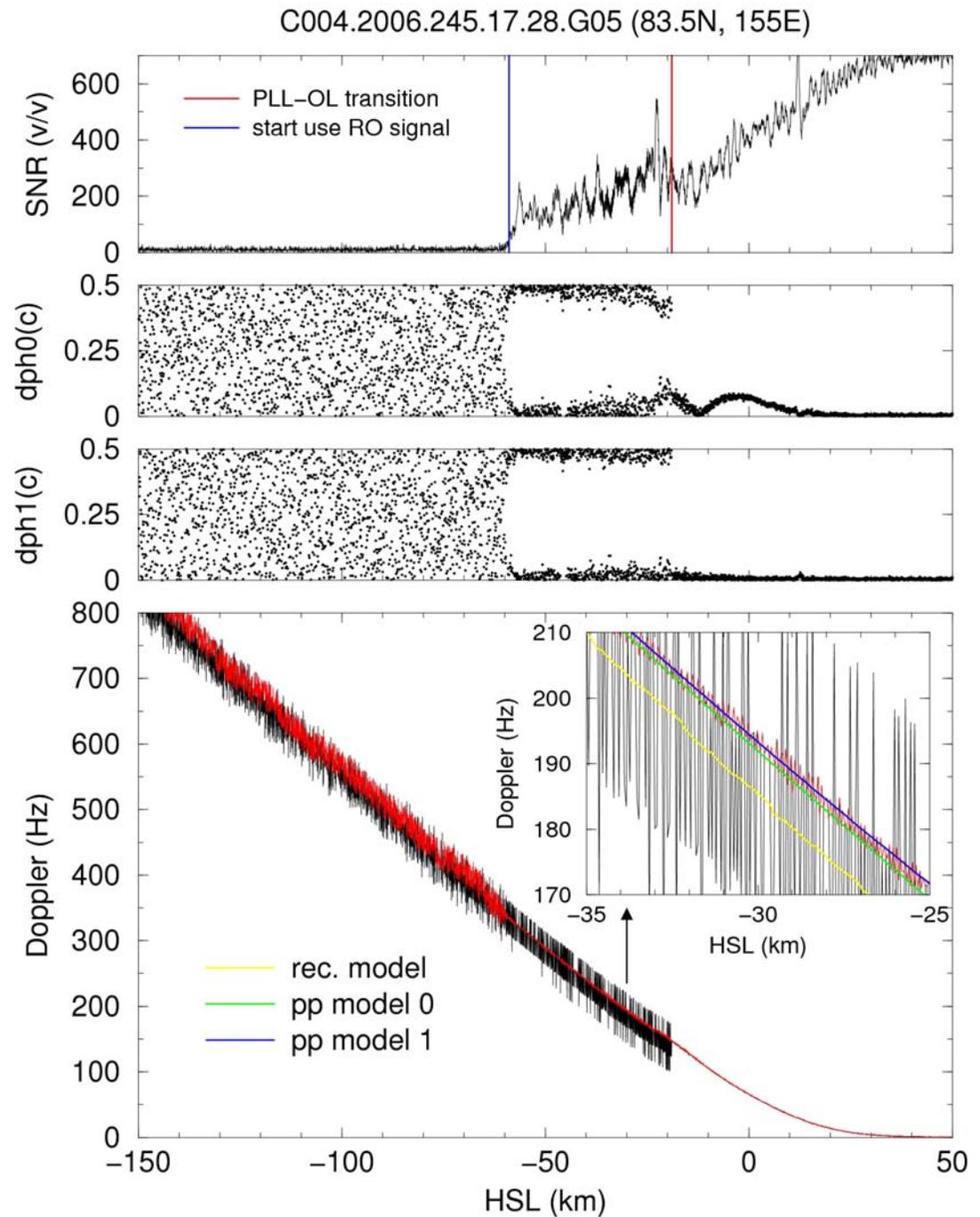
Noise level: 3 s averaged  
SNR at the end

Start use RO signal:  
1s-averaged SNR exceeds  
noise level by 50%

Phase lapse (atmosphere-  
-induced modulation +  
+ residual frequency +  
+ noise) < 1/4 cycle

NDM clearly distinguishable

internal NDM removal  
does **not** result in errors



## Low-latitude occultation

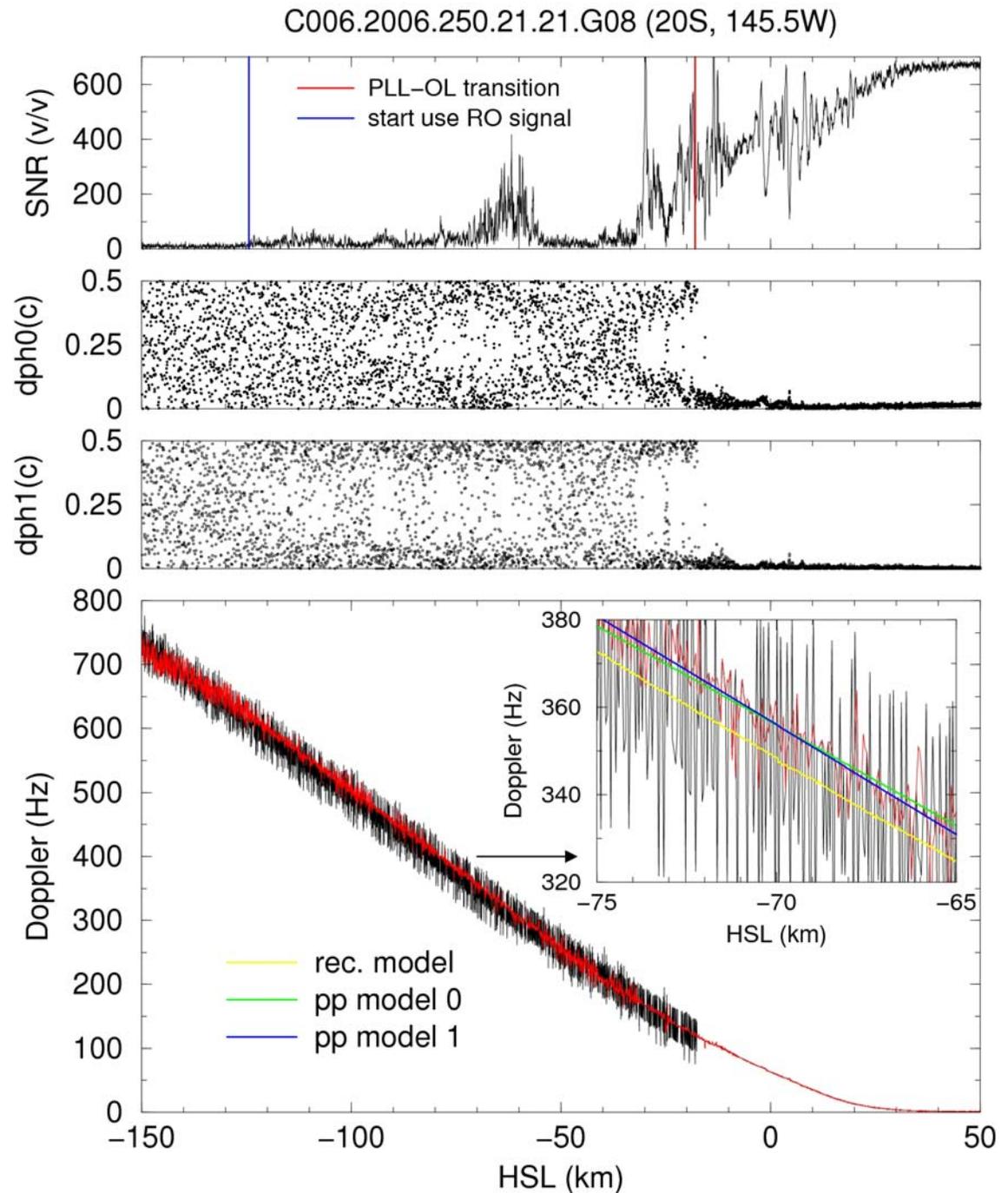
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Start use RO signal:  
1s-averaged SNR exceeds  
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Phase lapse (atmosphere-  
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+ noise)  $> 1/4$  cycle

NDM **not** clearly  
distinguishable

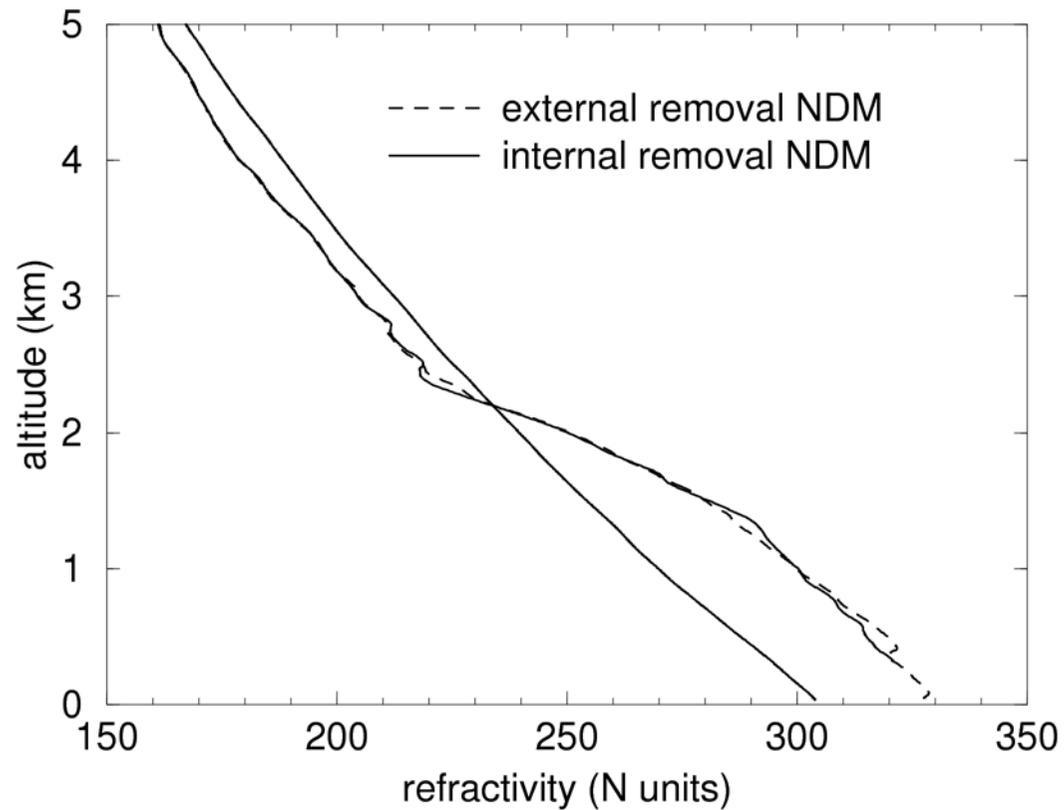
internal NDM removal  
**does** result in errors



# Refractivity profiles retrieved with internal and external removal of NDM for the two COSMIC occultations

High-latitude occultation - no difference

Low-latitude occultation - some difference



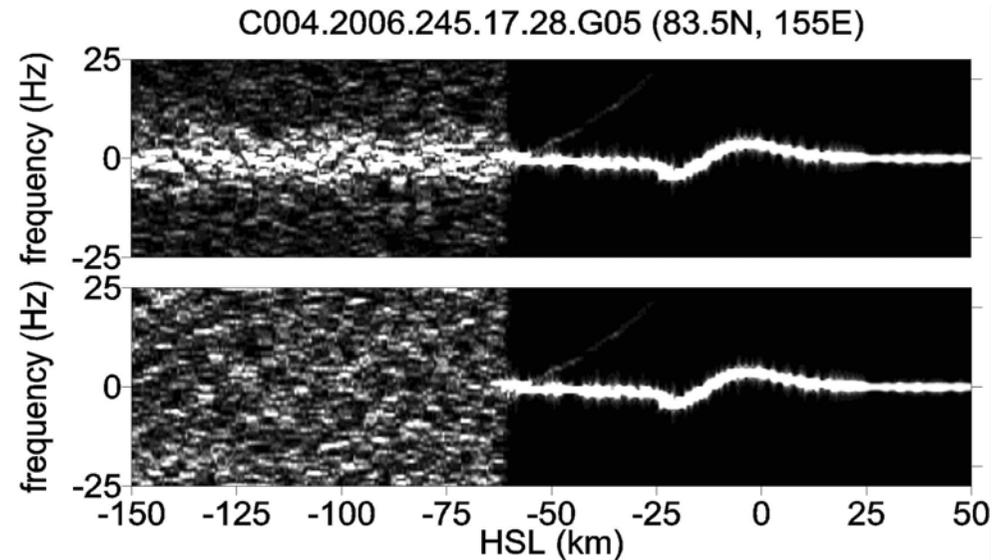
## From Model 0 to Model 1

Sliding spectrograms of RO signals. Down-conversion with Model 0.

### High-latitude occultation

- internal NDM removal

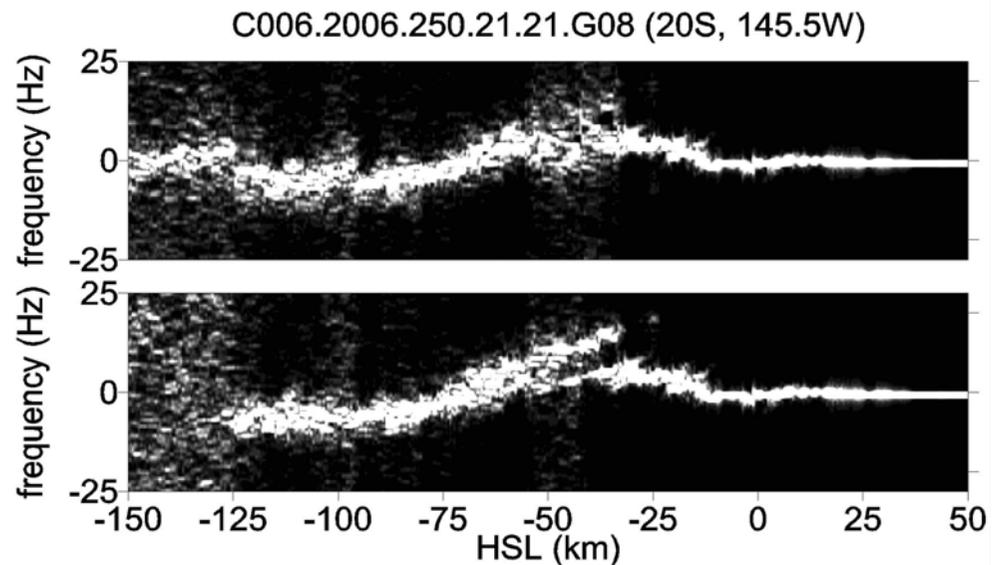
- external NDM removal



### Low-latitude occultation

- internal NDM removal

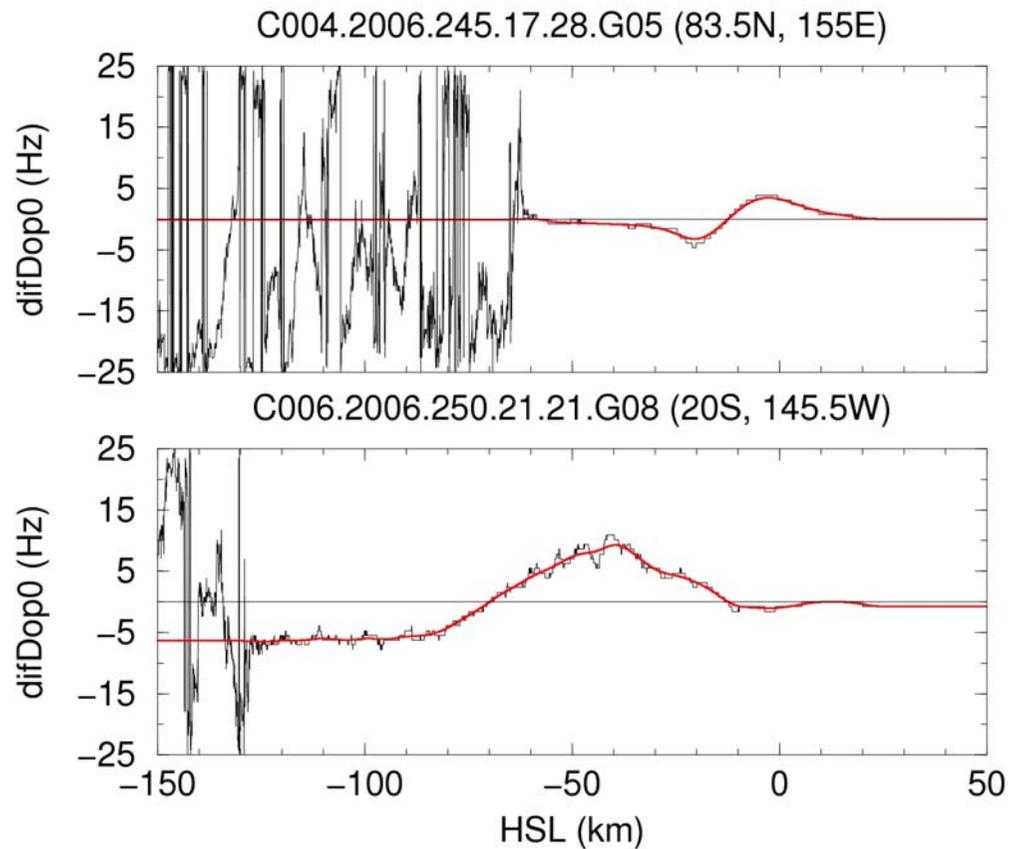
- external NDM removal  
(reveals multipath)



Frequency miss-modeling determined from the sliding spectrograms after down-conversion with **Model 0** can be used for adjustment of the frequency model

$$f_{\text{mod}}^1 = f_{\text{mod}}^0 + \Delta f$$

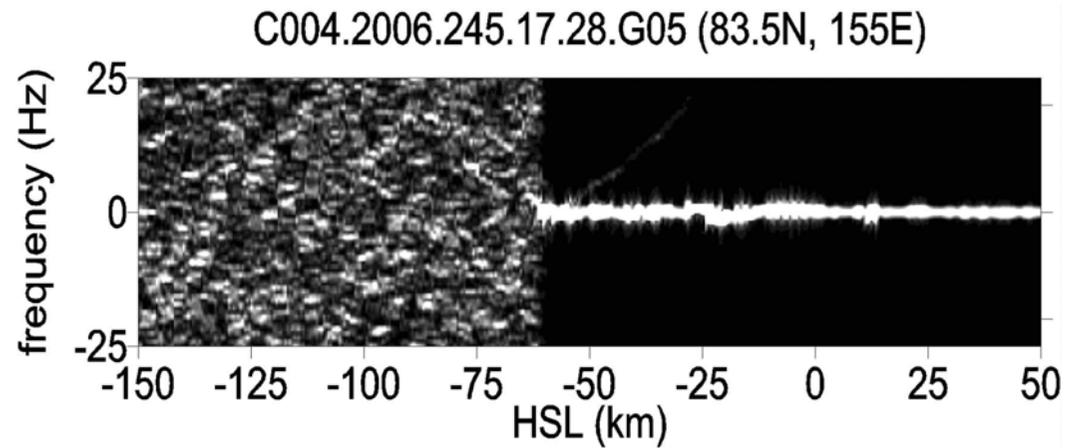
**High-latitude occultation**  
external NDM removal



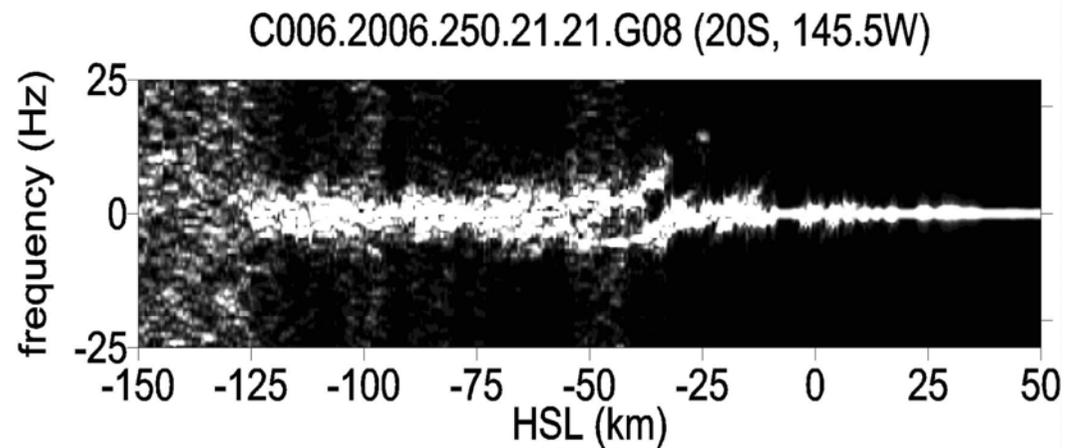
**Low-latitude occultation**  
external NDM removal

Sliding spectrograms of RO signals after down-conversion with **Model 1**

**High-latitude occultation**  
external NDM removal

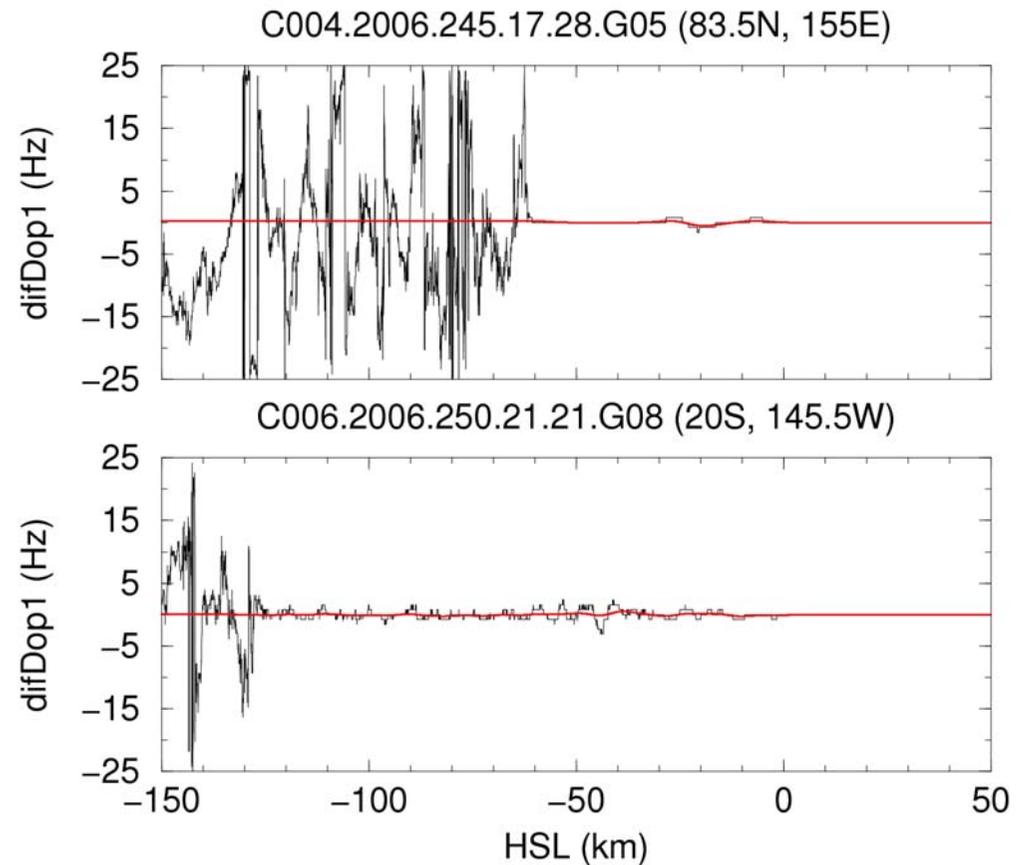


**Low-latitude occultation**  
external NDM removal



The residual frequency miss-modeling after down-conversion  
with **Model 1** is substantially reduced

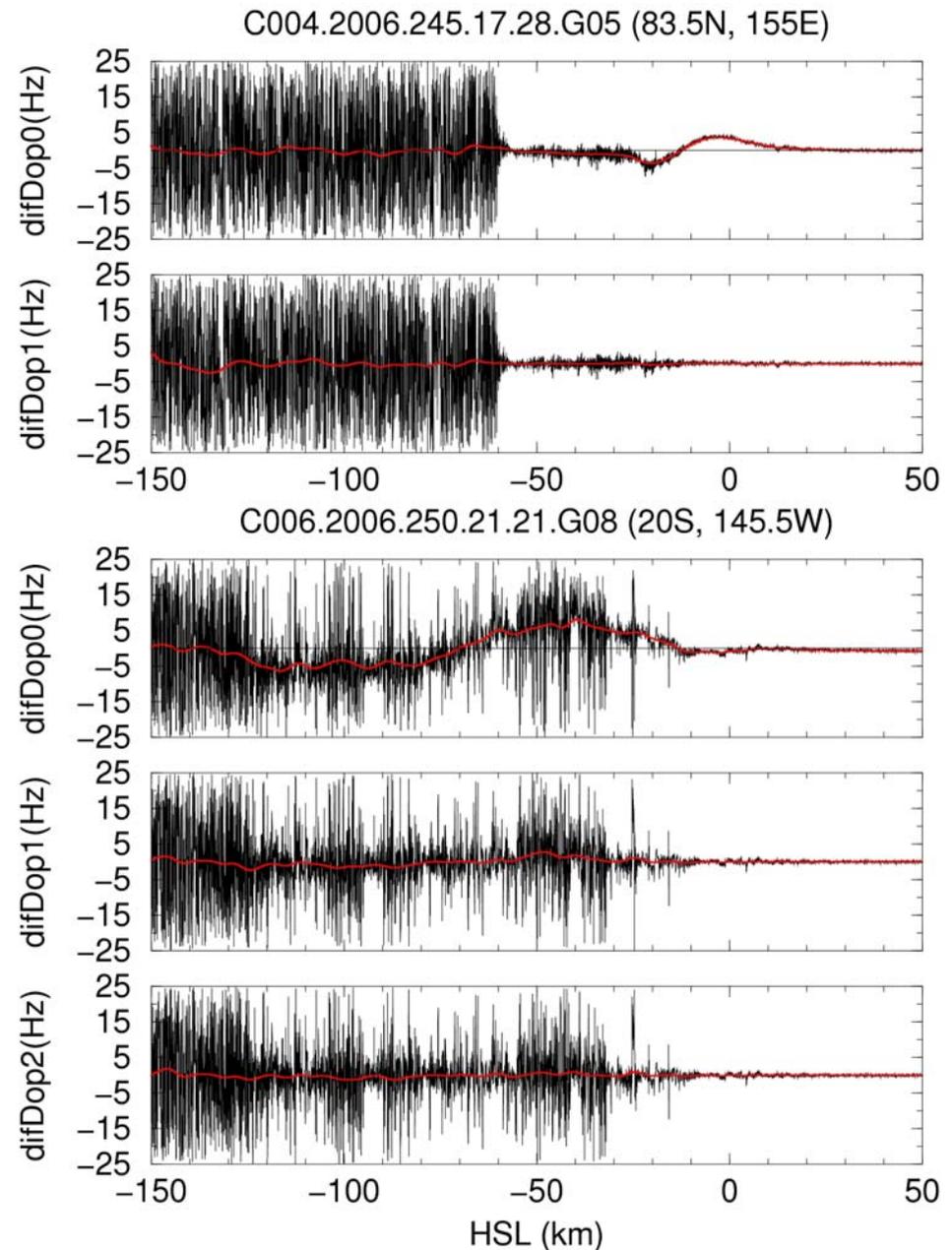
**High-latitude occultation**  
external NDM removal



**Low-latitude occultation**  
external NDM removal

Another method for determining frequency miss-modeling based on smoothed frequency after down-conversion with **Model 0**

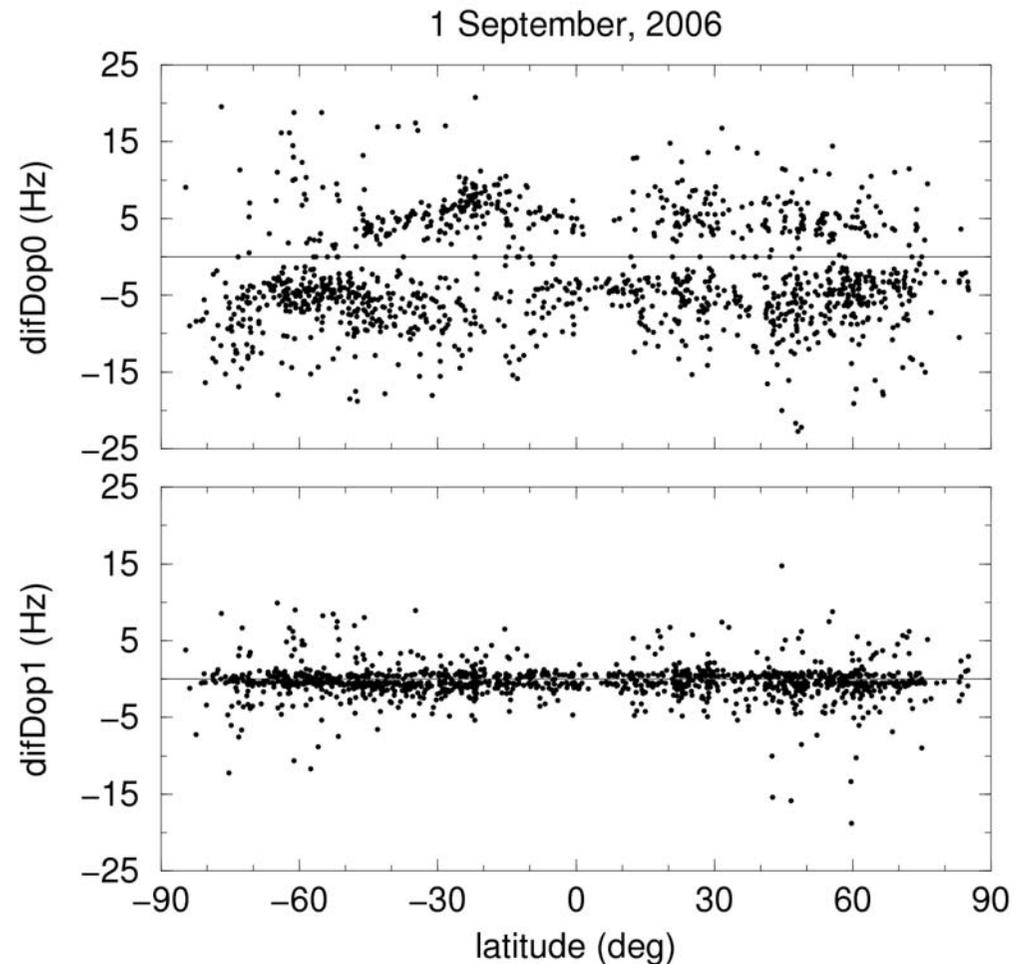
- computationally more simple than the spectrograms
- may need more iterations
- affected by cycle slips
- does not allow detection of the multipath



**Important:** when accuracy of first guess (**Model 0**) is within  $\pm 25$  Hz, the results of the post-processing after adjustment of the model do not depend on the first guess (**Model 0**).

Maximal frequency miss-modeling after down-conversion with **Model 0**

Maximal frequency miss-modeling after down-conversion with **Model 1**



## Statistical comparison of inverted COSMIC refractivities in LT in tropics (30S – 30N) for September 2006

10490 occultations:

7077 - externally recorded NDM bit sequences available

3413 - externally recorded NDM bit sequences not available

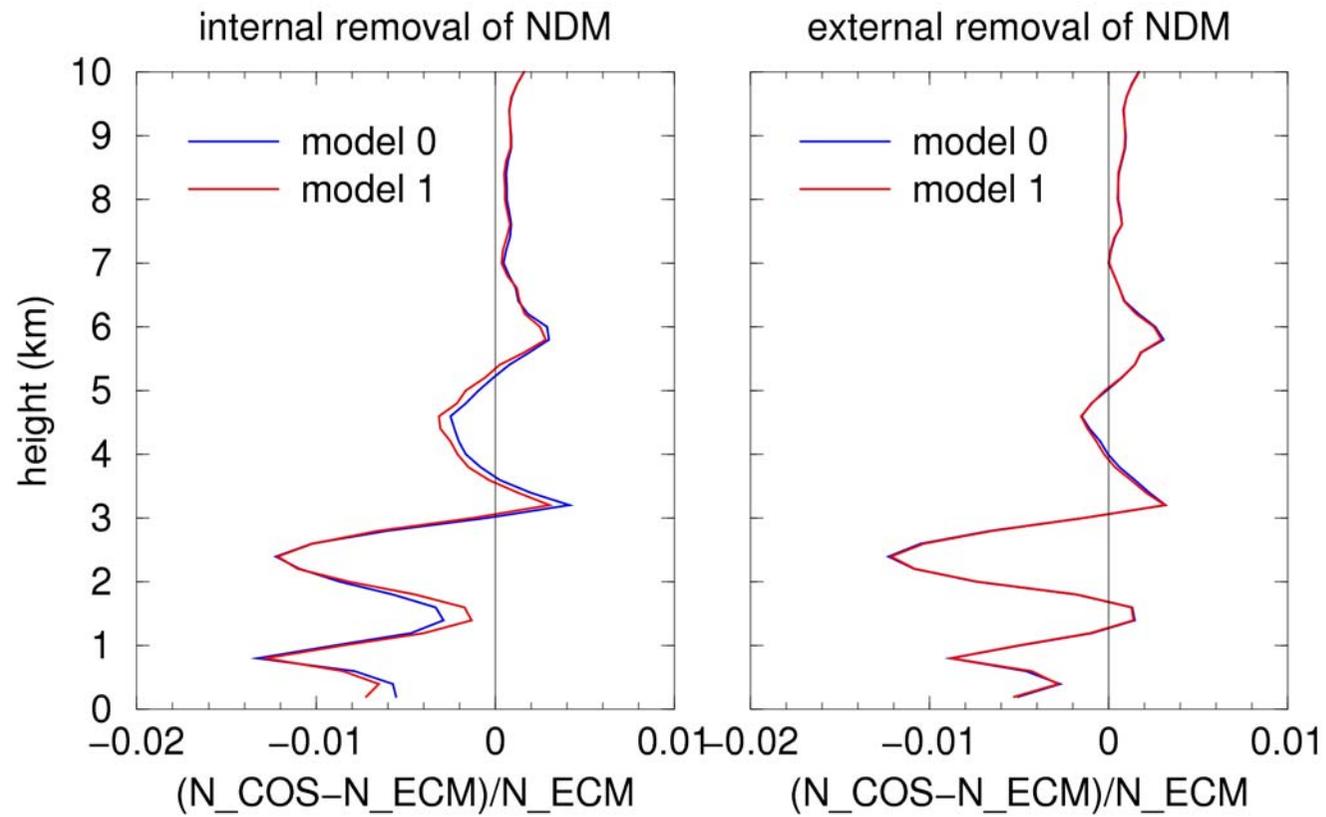
Only the occultations with available NDM were processed:

- with internal and external removal of NDM;
- with Model 0 and Model 1 for down-conversion

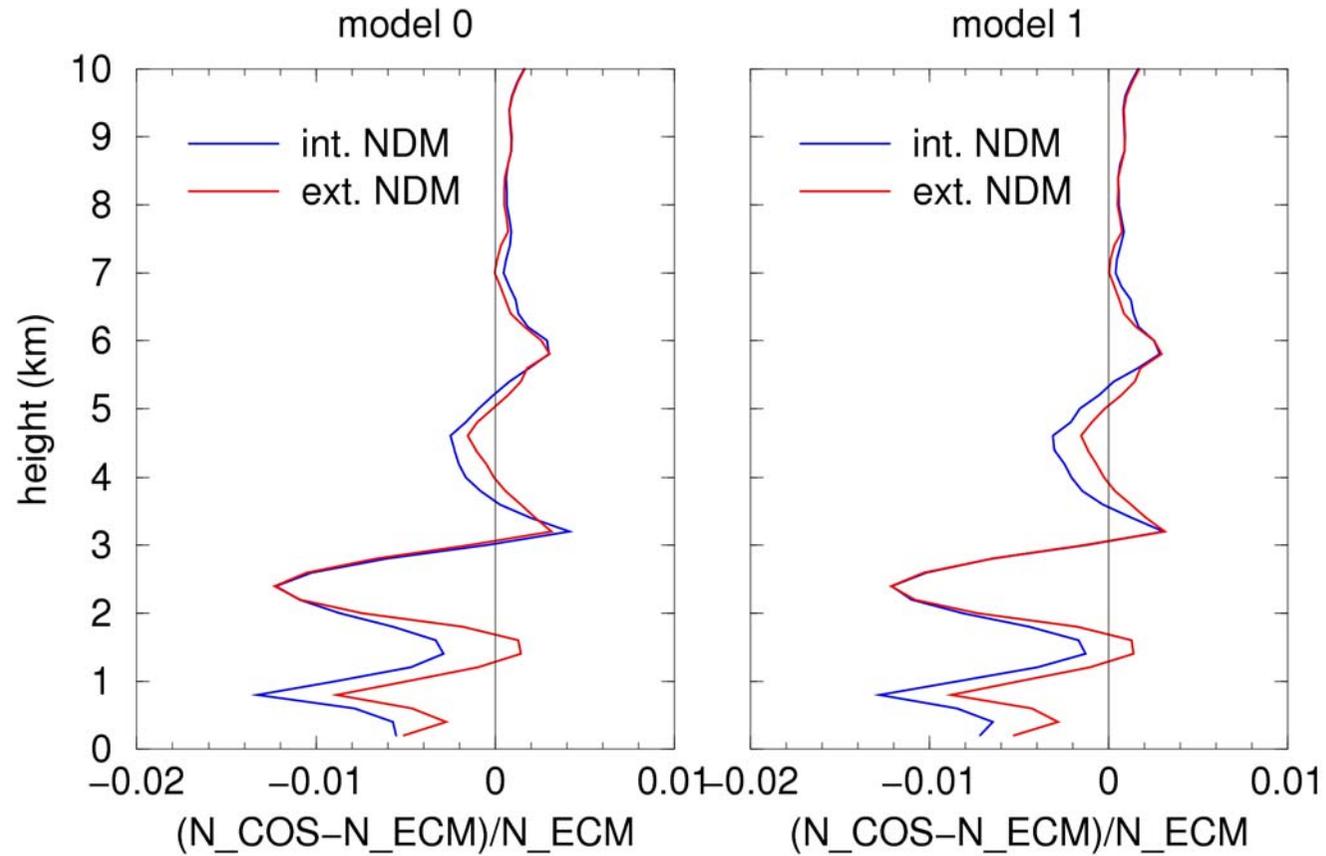
ECMWF 21-level analysis is used as the reference

- the differences between RO and ECMWF analysis (caused by different factors) are not considered;
- only the differences between different processing modes are considered;
- **important:** when a difference is found it is known which processing mode is responsible for the error

The first guess Model 0 (based on CIRA+Q) is sufficient when externally recorded NDM bit sequence is available

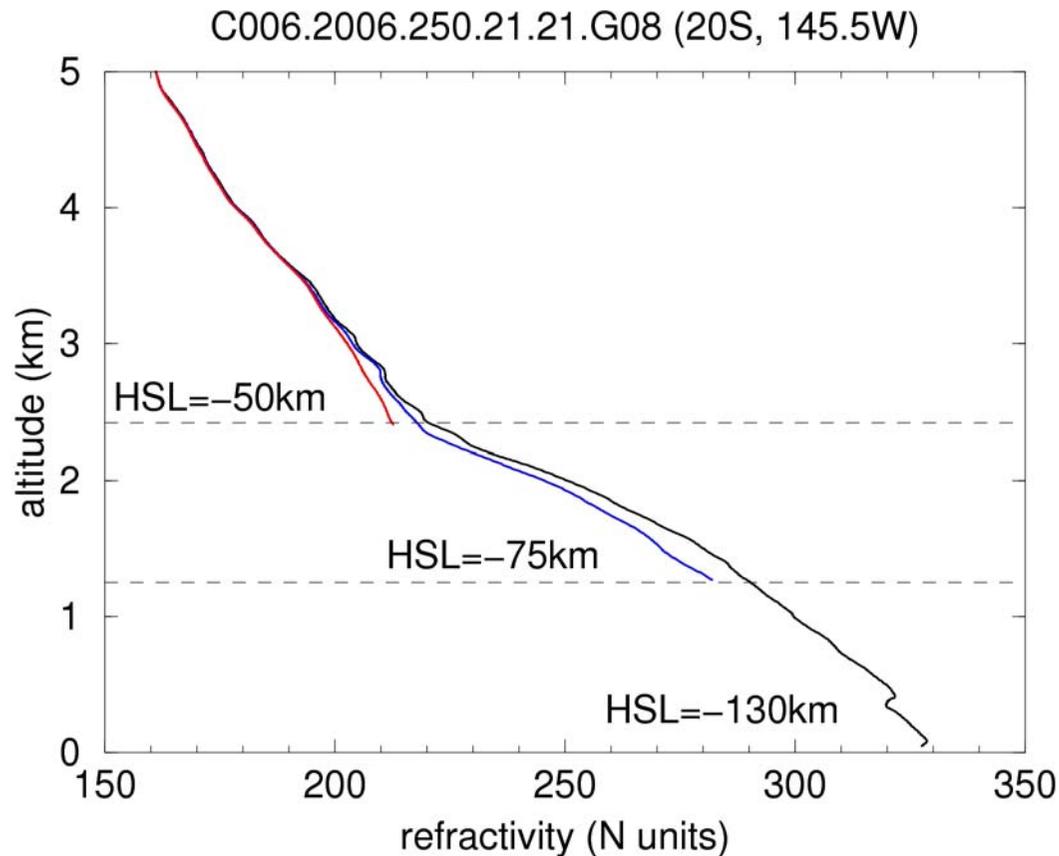


External removal of NDM results in reduction of the errors of the post-processing, even with the adjustment of the frequency model.

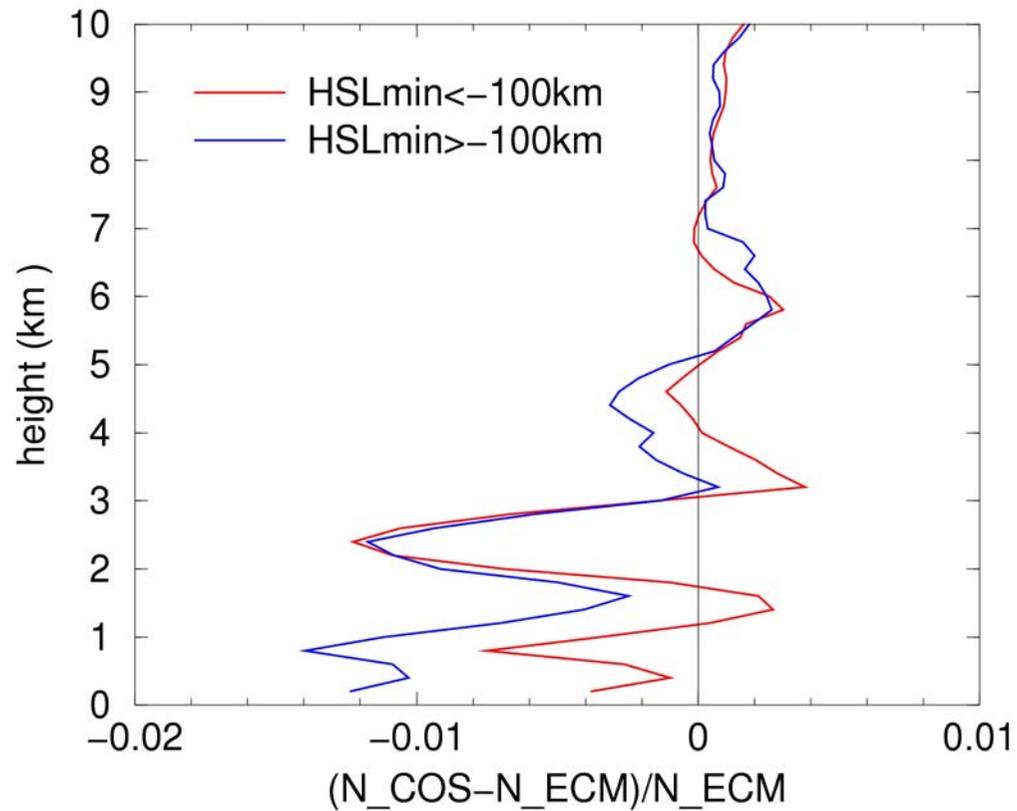


In tropical LT, RO signals propagate deep, at  $HSL < -100$  km due to diffraction by N-structures caused by WV. Due to the multipath, the relation between RO signal and retrieved profile becomes non-local. When RO signal is tracked not deep enough (truncated), this results in:

- loss of penetration (higher cut-off);
- inversion errors (bias) above the cut-off



Statistical validation of the effect of tracking depth in tropics (30S-30N)  
6634 occultations SNR exceeds noise level by 50% at HSL < -100 km  
3856 occultations SNR exceeds noise level by 50% at HSL > -100 km  
among possible reasons: loss of SNR due to errors of receiver models



# Conclusions

- receiver phase model is not used in the post-processing of OL signals
- adjustment of the frequency model (closing the loop) in the post-processing makes the results independent on the first guess when the first guess model is accurate within +/-25 Hz
- with the external removal of NDM, the phase model based on CIRA+Q is sufficient for the post-processing of all occultations
- with the internal removal of NDM the phase model based on CIRA+Q is sufficient for high-latitude occultations and sufficient as the first guess with the following adjustment for tropical occultations
- all COSMIC occultations are processed with the adjustment of the phase model (NDM bit sequences not available for all occultations)
- reduction of the errors due to insufficient tracking depth in tropics needs better accuracy of receiver models