

Investigation on Formosat-3 /COSMIC retrieval data quality

Yuei-An Liou^{*}, Shiang-Kun Yan, Li-Ru Chen, and
Cheng Yung Huang

Center for Space and Remote Sensing Research
National Central University

2010.09.09 OPAC2010

Motivation

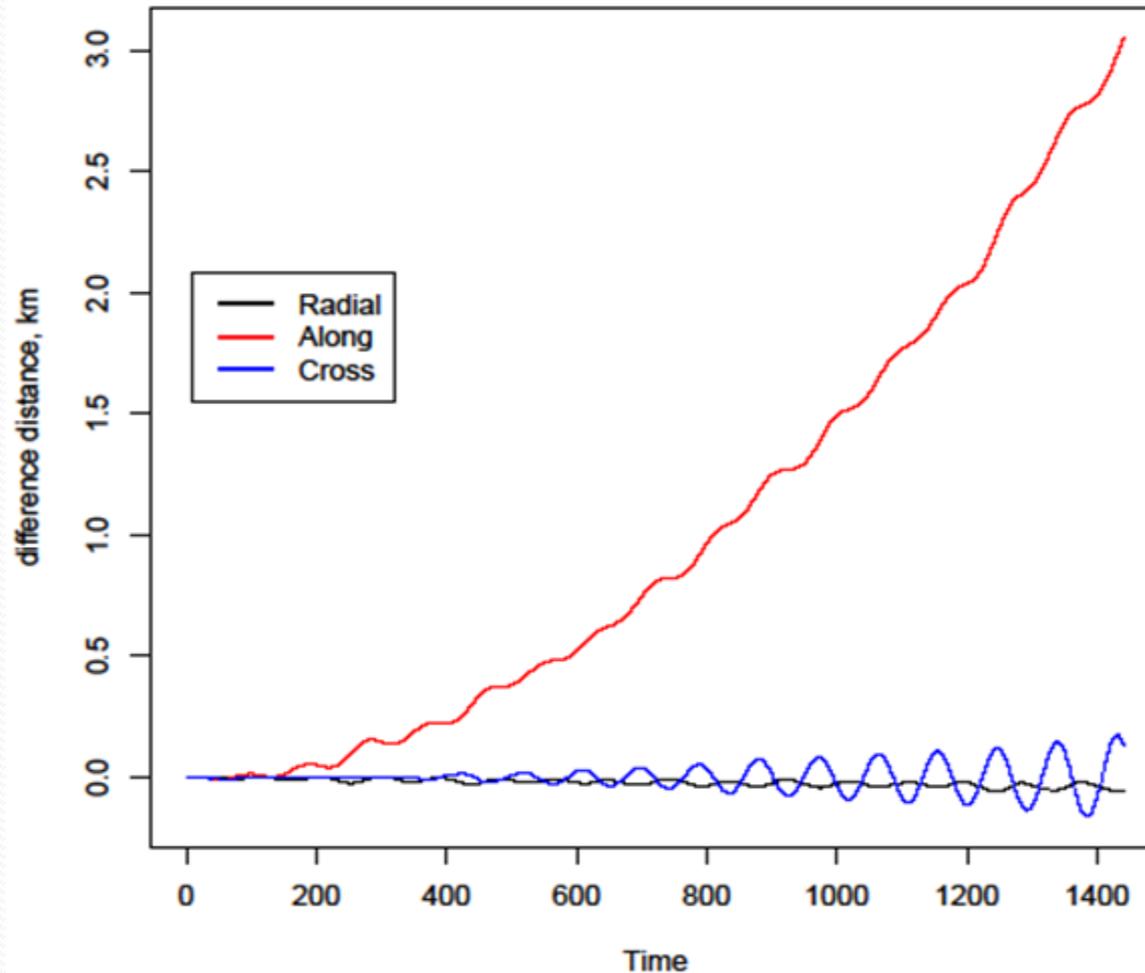
- Initial purpose:
 - Calculate neutral density (ρC_D) from FS₃/COSMIC orbital change while the altitude park around 500 km.
 - N-body and Solar radiation pressure(specular reflection) are considered
 - Albedo, IR, Solid Tide, Ocean Tide, Lift and Side are not considered.
- The drag just could be get while the resolution duration is longer than 3 or even more orbits.
- GOAL: get the higher time resolution for scientific research.

Orbital difference (SP3-Ref orbit)

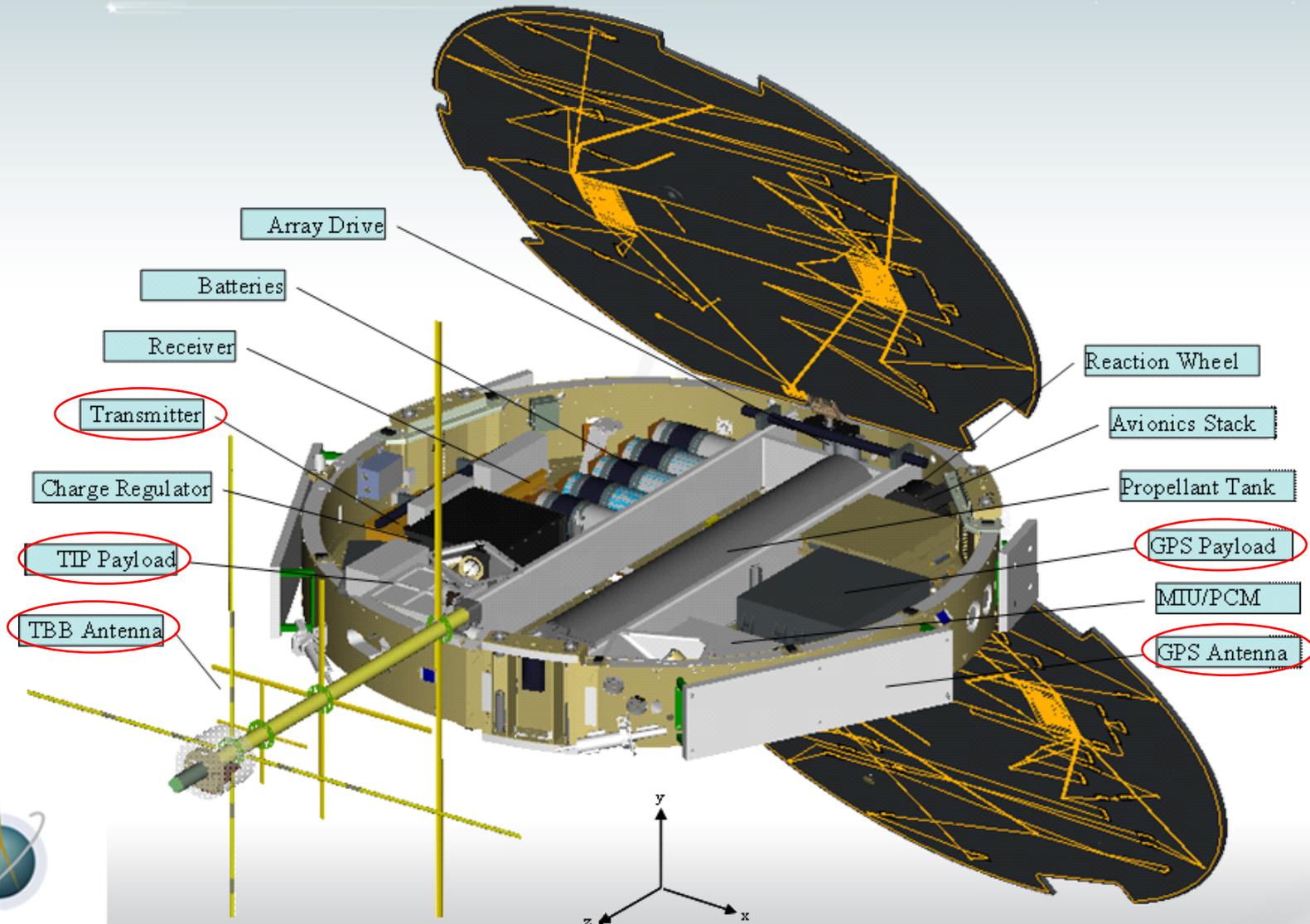
Ref orbit :

gravity (ICGEM 70order),

DE405 (Sun and moon and 8 planets)



Spacecraft Overview



Solar radiation pressure (David A. Vallado, 2007)

Solar Radiation Pressure

$$\mathbf{a}_{SR} = -\sum_{i=1} \frac{P_{SR} A_{\odot} \cos(\phi_{inc})}{mc} \left\{ 2\left(\frac{C_{Rd_i}}{3} + C_{Rs_i} \cos(\phi_{inc})\right)\mathbf{n} + (1 - C_{Rs_i})\mathbf{s} \right\}$$

C_{Ra} : absorption

C_{Rs} : specular reflection

C_{Rd} : diffuse reflection

$C_{Ra} + C_{Rs} + C_{Rd} = 1$

$P_{SR} = 1367 \text{ W/m}^2 = 4.57\text{E}+6 \text{ N/m}^2$

Atmospheric Drag

$$\mathbf{a}_{drag} = -\frac{1}{2} \frac{C_D A}{m} \rho v_{rel}^2 \frac{\mathbf{v}_{rel}}{|\mathbf{v}_{rel}|}$$

C_D : drag coefficient

Albedo and IR is similar to Solar Radiation.

Parameters

- Orbital data : LeoOrb
- Satellite attitude : LeoAtt
- Neutral Density: JB2008 (Developed by Air Force)
- Solar panel

$$C_{Ra} = 0$$

$$C_{Rs} = 0.5$$

$$C_{Rd} = 0.5$$

$$C_{Rd} = 2.2$$

$$C_{Lift} = 2.2$$

$$C_{Side} = 2.2$$

$$P_{SR} = 1367 \text{ W/m}^2 = 4.57\text{E}+6 \text{ N/m}^2$$

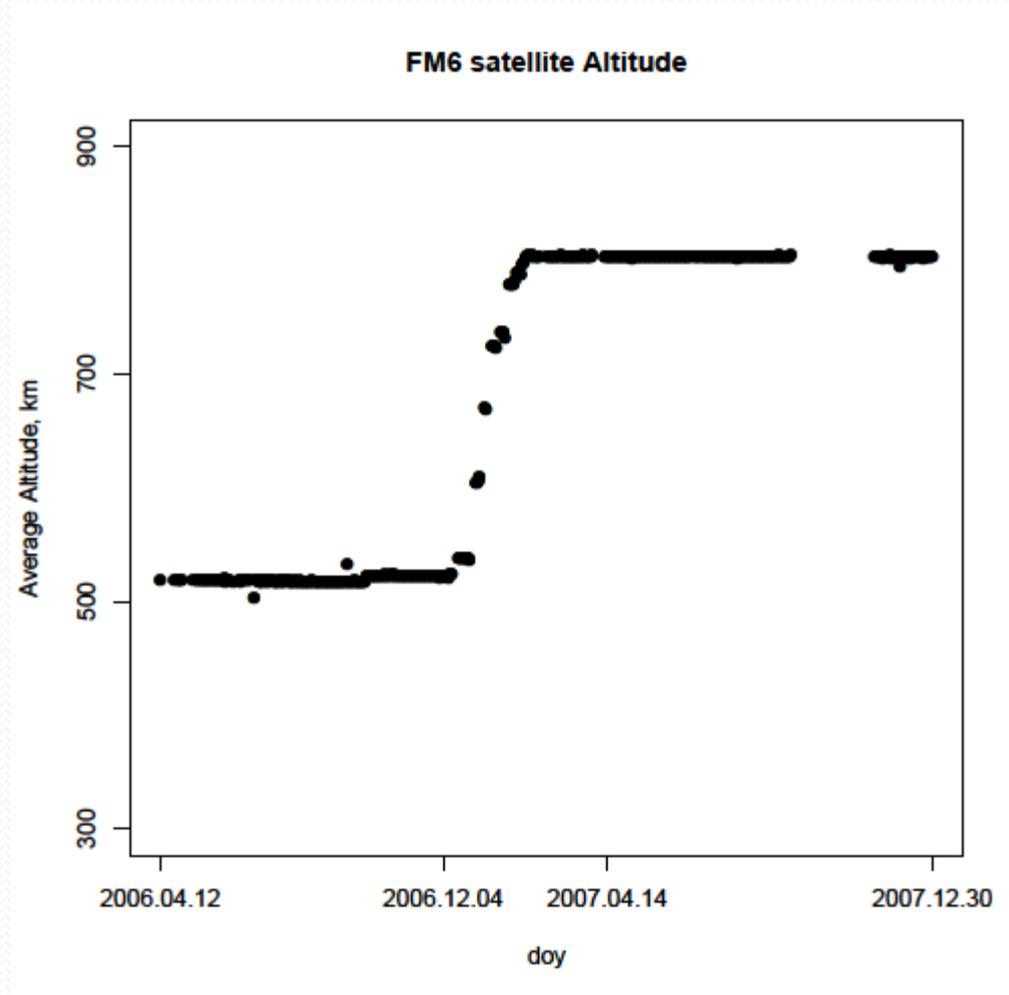
$$P_{Albedo} = 0.3P_{SR} = 410 \text{ W/m}^2$$

$$P_{IR} = 237 \text{ W/m}^2$$

- Cylinder body

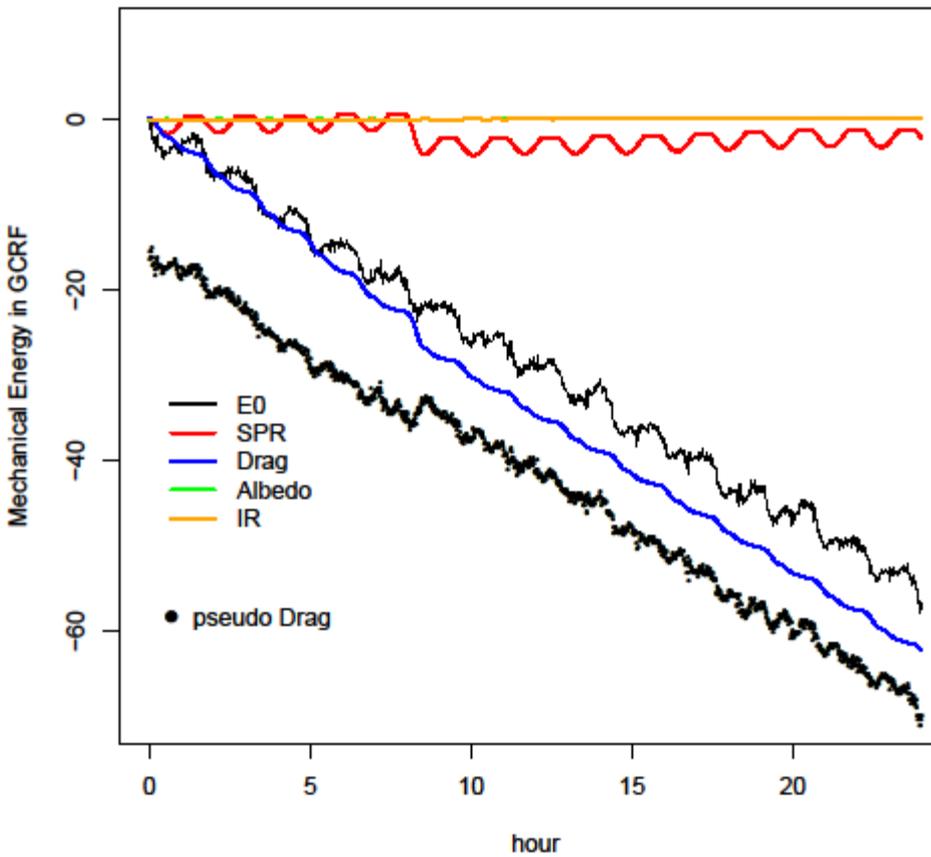
$$C_{Rd} = 1.17$$

FM6 Altitude

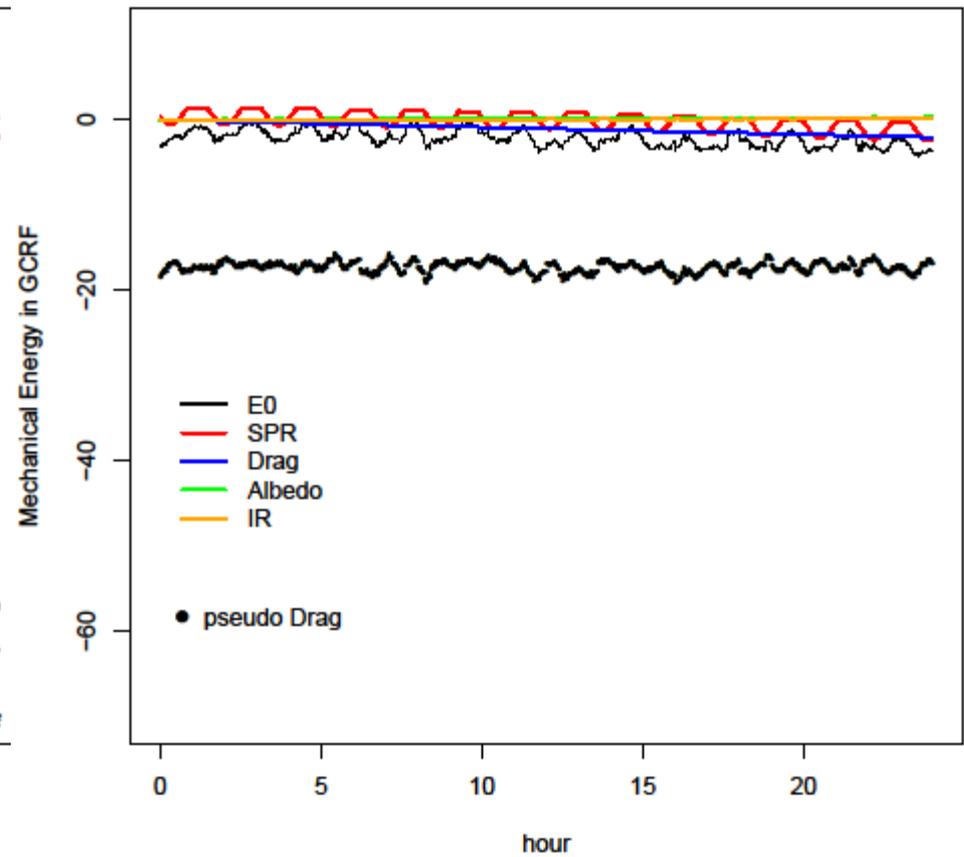


$$E = -\frac{\mu}{2a}$$

leoOrb_2006.338.006

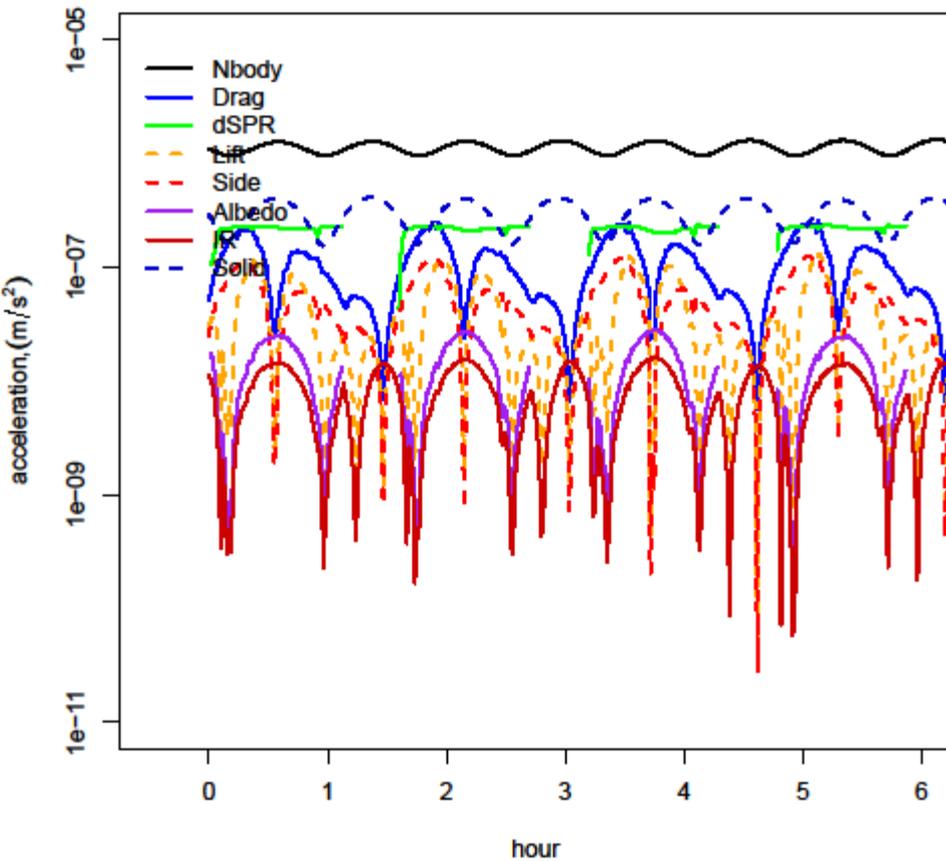


leoOrb_2007.103.006

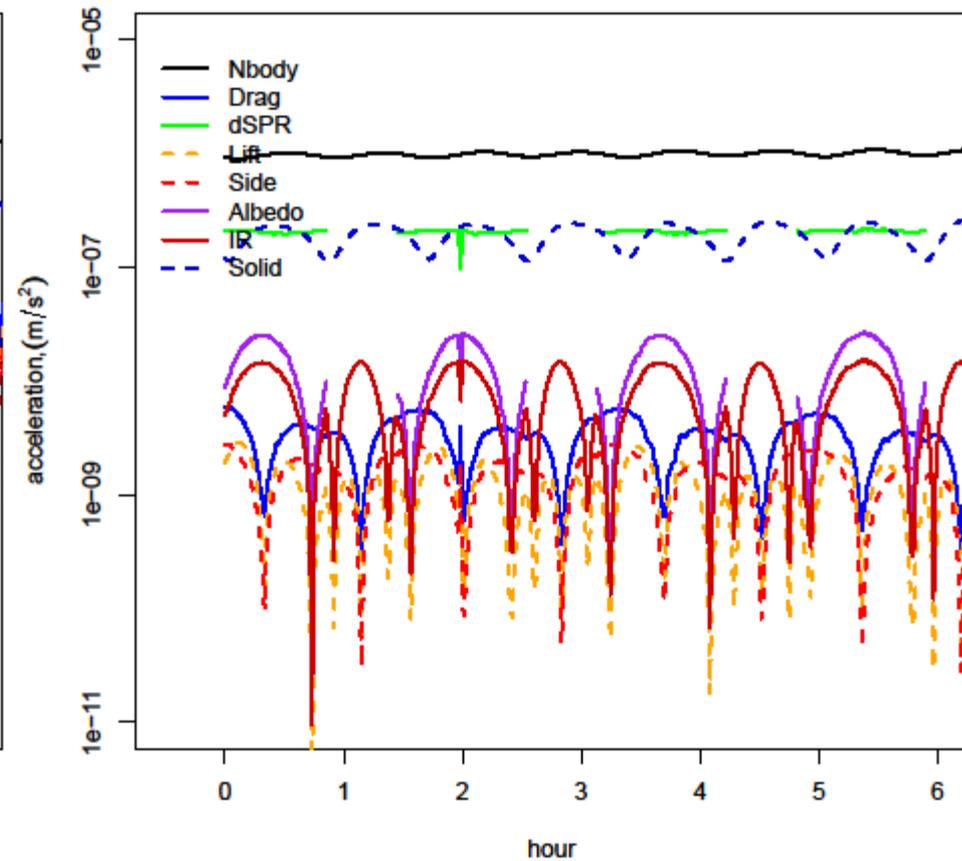


Accerleration analysis

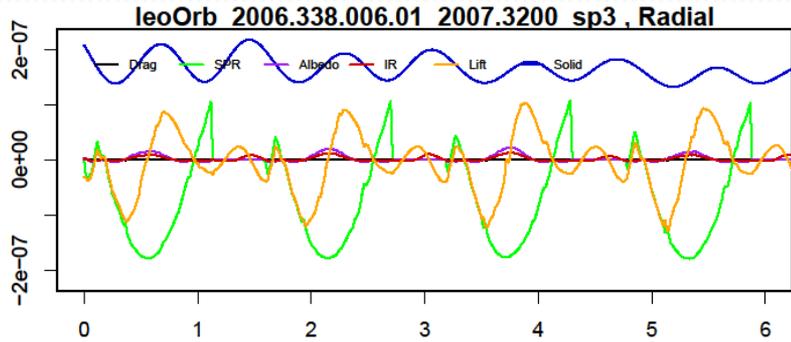
leoOrb_2006.338.006.01_2007.3200_sp3



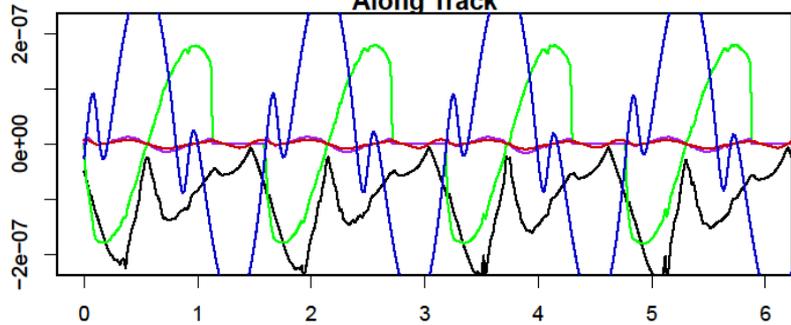
leoOrb_2007.103.006.01_2007.3200_sp3



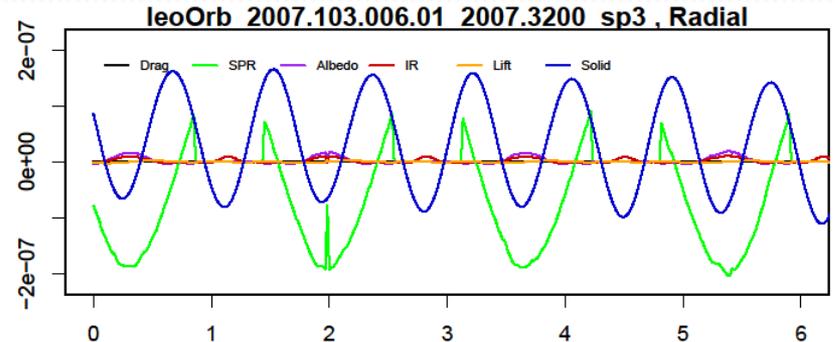
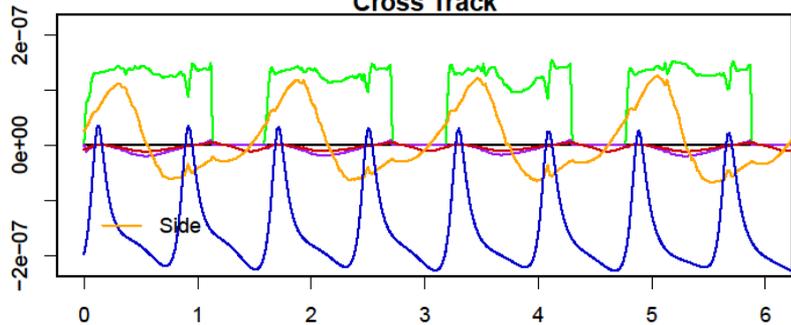
Accerleration analysis in RSW



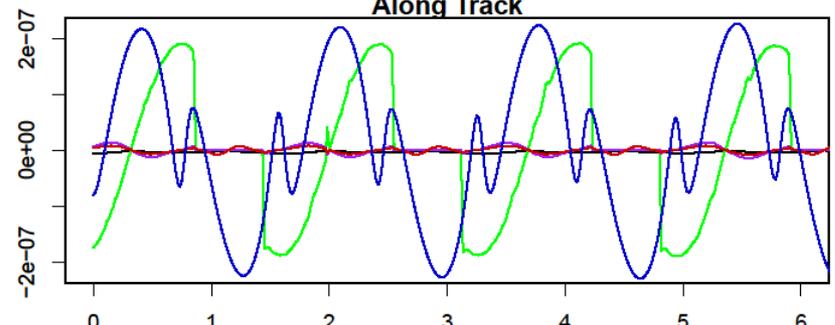
Along Track



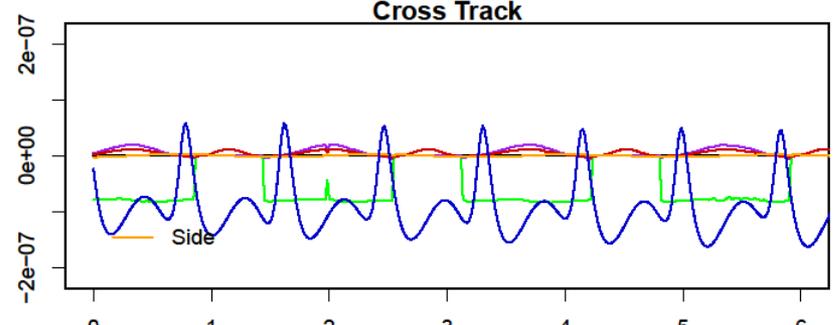
Cross Track



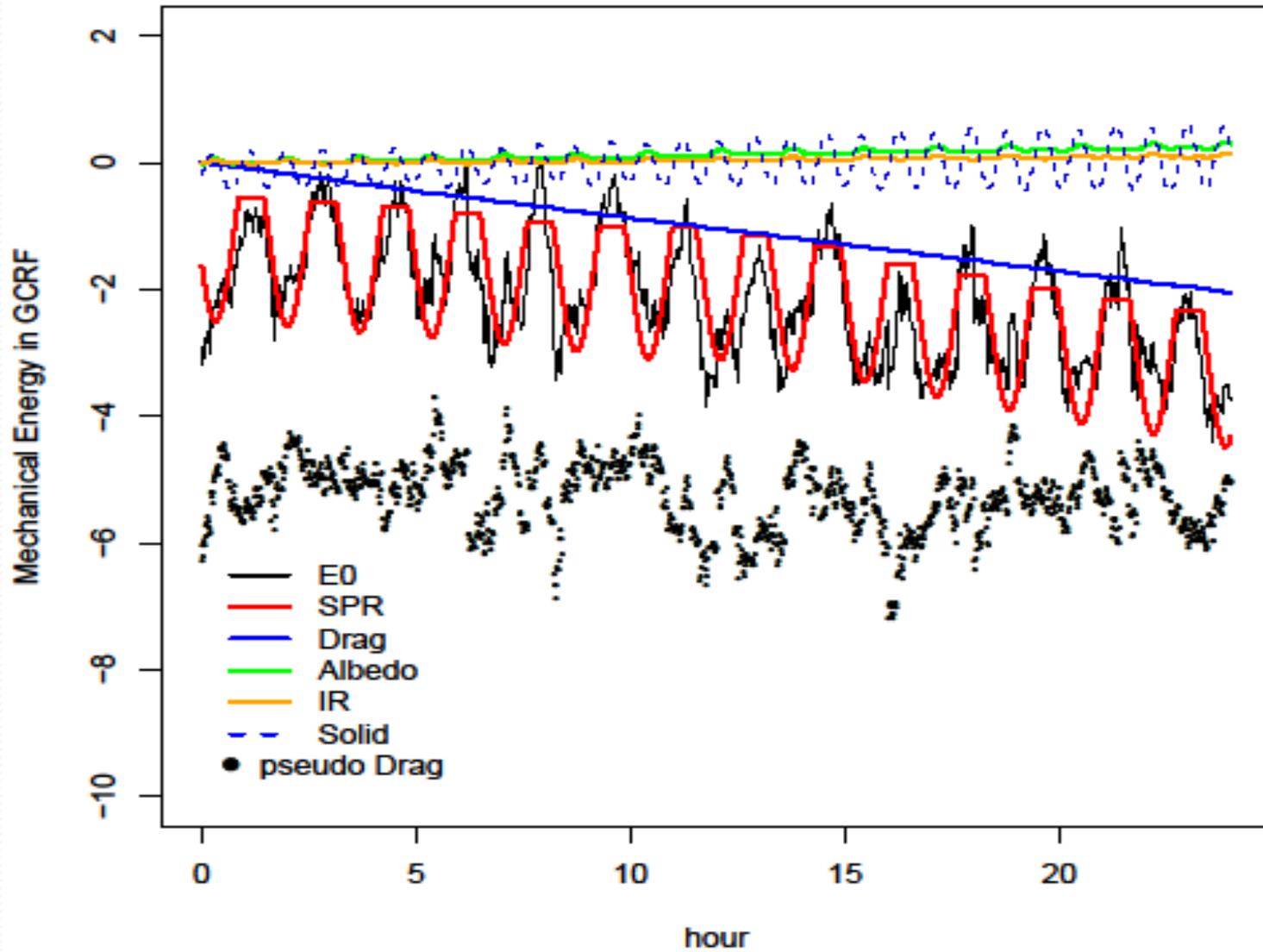
Along Track

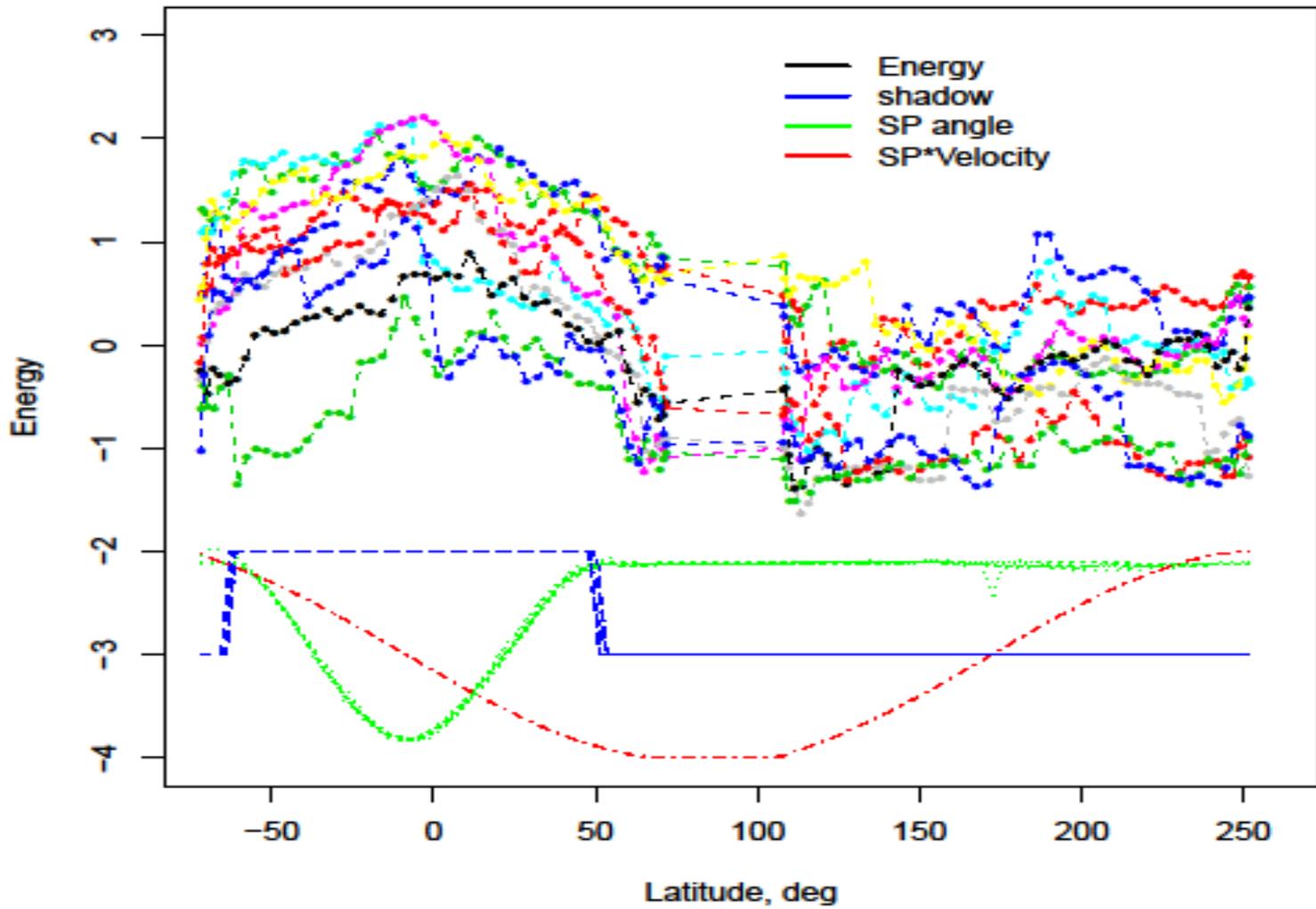


Cross Track



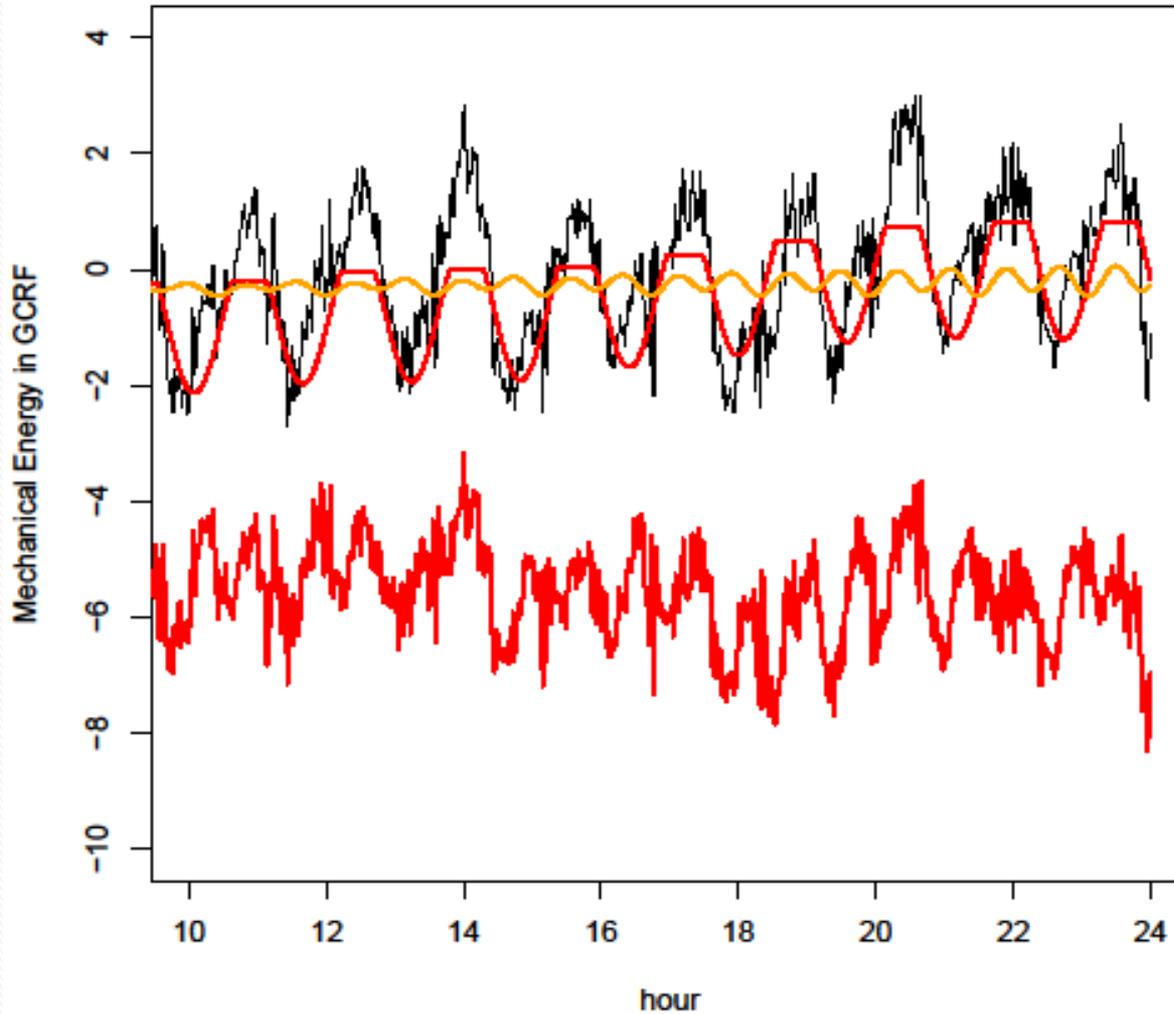
leoOrb_2007.103.006



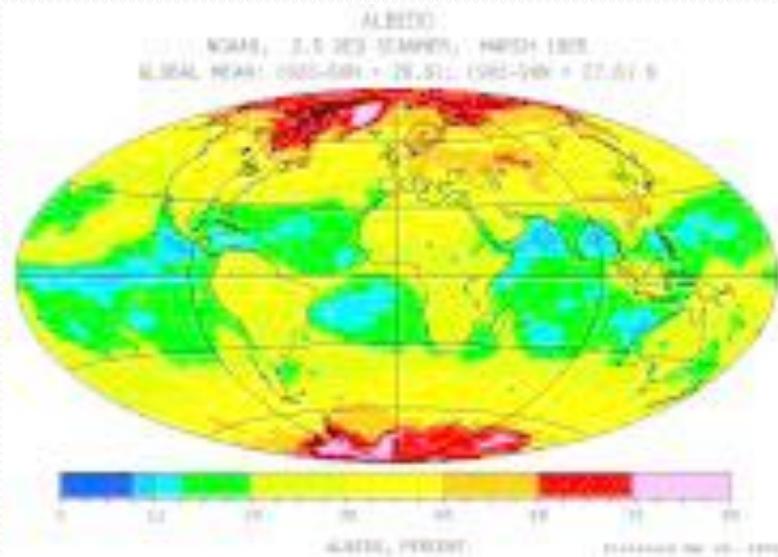


Orbit analysis around 500 km

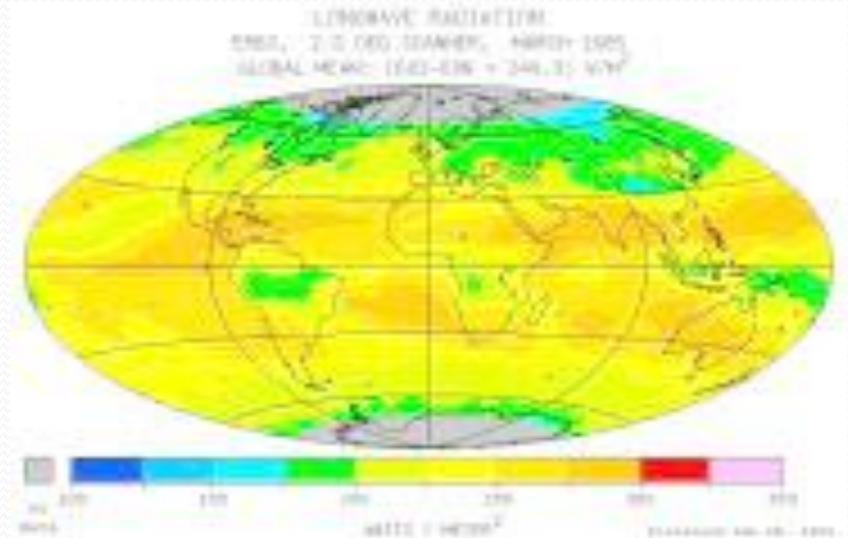
leoOrb_2006.338.006



Albedo and IR emission



*Planetary albedo map composed from NOAA-9 data
(based on observations in March 1985)*



*Earth IR re-radiation map composed from ERBE data
(based on observations in March 1985)*

Conclusions

→ Bigger Amplitude fluctuation seem could be removed by finding the correct solar panel reflection, diffuse and absorption coefficients from orbit data around 800 km .

→ What we will do in future

1. Observed IR and Albedo data should be used,
2. Higher resolution orbital data maybe used,
3. Ocean Tide.

→ Challenge

1. Small amplitude fluctuation
2. Where Error might come from:
 - i. Integrator,
 - ii. SP3 orbit,
 - iii. Wind around 500 km.
3. Unconsidered perturbation,